LA-UR-13-24322

Approved for public release; distribution is unlimited.

Title: The Science Program at the Los Alamos Ultracold
Neutron Source

Author(s): Saunders, Alexander

Intended for: PSI seminar June 2013
Report

Issued: 2013-06-13

Disclaimer:

Los Alamos National Laboratory, an affirmative action/equal opportunity employer, is operated by the Los Alamos National Security, LLC for the National Nuclear Security Administration of the U.S. Department of Energy under contract DE-AC52-06NA25396. By approving this article, the publisher recognizes that the U.S. Government retains nonexclusive, royalty-free license to publish or reproduce the published form of this contribution, or to allow others to do so, for U.S. Government purposes. Los Alamos National Laboratory requests that the publisher identify this article as work performed under the auspices of the U.S. Department of Energy. Los Alamos National Laboratory strongly supports academic freedom and a researcher's right to publish; as an institution, however, the Laboratory does not endorse the viewpoint of a publication or guarantee its technical correctness.

The Science Program at the Los Alamos Ultracold Neutron Source

Alexander Saunders
Los Alamos National Laboratory

PSI Seminar, 2013

Outline

- Introduction
 - Motivation of neutron β decay measurements
 - UCNA current status and plans
 - A new UCN lifetime measurement
 - UCNB plans and goals
 - The LANSCE UCN source
 - Summary
-
- Special Thanks to Mark Makela, Alejandro Garcia, Brad Plaster, Albert Young, Chen-Yu Liu, Dirk Dubbers, and Kevin Hickerson

Neutron Decay Parameters

- Semi-leptonic decay
 - Lifetime 880 s
 - Endpoint energy 782 keV
- Just two free parameters in SM
 - CKM mixing matrix element
 - Ratio of weak coupling constants
 - Uncertainty comes from radiative corrections

$$n \rightarrow p + e^- + \bar{\nu}_e$$

$$\tau_n = \frac{4908.7 \pm 1.9 \text{ s}}{|V_{ud}|^2 (1 + 3\lambda^2)}$$

$$\lambda = g_A / g_V$$

Neutron β decay and V_{ud}

Angular correlations in polarized neutron decay (Jackson *et al* '57)

$$d\Gamma = d\Gamma_0 \times \left[1 + a \frac{\vec{p}_e \cdot \vec{p}_\nu}{E_e E_\nu} + b \frac{m_e}{E_e} + \langle \vec{\sigma}_n \rangle \cdot \left(A \frac{\vec{p}_e}{E_e} + B \frac{\vec{p}_\nu}{E_\nu} + D \frac{\vec{p}_e}{E_e} \times \frac{\vec{p}_\nu}{E_\nu} \right) \right]$$

$$a = \frac{1 - |\lambda|^2}{1 + 3|\lambda|^2}, \quad A = -2 \frac{|\lambda|^2 + \text{Re}(\lambda)}{1 + 3|\lambda|^2}, \quad B = 2 \frac{|\lambda|^2 - \text{Re}(\lambda)}{1 + 3|\lambda|^2}, \quad b_n = \frac{|b_F| - 3\lambda |b_{GT}|}{1 + 3\lambda^2}$$

$$\tau^{-1} = K f^R (G_V^2 + 3G_A^2) \quad B = B_0 + B_1 \frac{m_e}{E_e}$$

$$\lambda \equiv \frac{G_A}{G_V}$$

In Standard Model:

$$B_1 = 0 \quad b = 0$$

Neutron beta decay can inform many areas of physics

- Many reactions share the same Feynman diagram as neutron beta decay

Primordial element formation
(^2H , ^3He , ^4He , ^7Li , ...)

Solar cycle

Neutron star formation

Pion decay

Neutrino detectors

Neutrino forward scattering

W and Z production

Effects of $\lambda = G_A/G_V$

Where does the proton get its spin?

Some comes from the quarks, but how much and which quarks?

$G_A^3 = \Delta u - \Delta d$ Axial-vector weak charge Neutrons *et al.*

$G_A^8 = \Delta u + \Delta d - 2\Delta s$ SU(3) octet weak charge hyperon decay

$G_A^0 = \Delta u + \Delta d + \Delta s$ Flavor singlet weak charge DIS

Bass, 2005

Solar fusion rates and the Standard Solar Model

$$r_{pp} \propto \langle \sigma v \rangle_{pp} \propto \int_0^\infty E \sigma(E) e^{\left(\frac{-E}{kT}\right)} dE$$

$$\sigma(E) = \frac{S(E)}{E} e^{\frac{-2\pi\alpha}{v(E)}} \quad S_{pp} \propto \left(\frac{G_A}{G_V} \right)^2$$

Adelberger 2010

Effects of $\lambda = G_A/G_V$

- g_A is a fundamental parameter for the charged weak current of the nucleon and defines a frontier for our understanding of this interaction
- Lattice calculations of g_A are becoming an interesting target for high precision lattice calculations (Yamazaki et al., 2008)
- Reactor antineutrino anomaly: g_A affects antineutrino cross sections needed for that work (Mention et al., ArXiv: 1011.2755v4)

Neutron Decay and Unitarity

$$\begin{pmatrix} d_w \\ s_w \\ b_w \end{pmatrix} = \begin{pmatrix} V_{ud} & V_{us} & V_{ub} \\ V_{cd} & V_{cs} & V_{cb} \\ V_{td} & V_{ts} & V_{tb} \end{pmatrix} \begin{pmatrix} d_m \\ s_m \\ b_m \end{pmatrix}$$
$$\begin{pmatrix} d_w \\ s_w \\ b_w \end{pmatrix} = \begin{pmatrix} 0.975 & 0.22 & 0.005 \\ 0.22 & 0.97 & 0.04 \\ 0.005 & 0.04 & 0.99 \end{pmatrix} \begin{pmatrix} d_m \\ s_m \\ b_m \end{pmatrix}$$

Unitarity, eg. $|V_{ud}|^2 + |V_{us}|^2 + |V_{ub}|^2 = 1$, (or lack thereof) of CKM matrix tests existence of further quark generations and possible new physics (eg. Supersymmetry)

Motivation for A measurement: independent determination of V_{ud}

Motivation for b and B measurement: BSM Physics

- ε_s and ε_t are non-(V-A) contributions to the effective Hamiltonian

Search for new physics

- A single parameter yields λ , multiple measurements yield V_{ud} and beyond
- CKM unitarity $|V_{ud}|^2 + |V_{us}|^2 + |V_{ub}|^2 = 1$
 - Do neutrons and superallowed beta decays agree?
- Search for right-handed currents (250 GeV limit from n decay)
- Scalar and tensor couplings from B and b
 - Cirigliano 2012

Holeczek *et al.*, arxiv 1303.5295 (2013)

Phenomenology: preliminary study

- Constraints on ε_S and ε_T from proposed experiments vs other measurements ($0^+ \rightarrow 0^+$; $\pi \rightarrow e \nu \gamma$ and collider not competitive)

- Already 10^{-3} measurements probe multi-TeV scale; stringent constraint on existence of scalar and tensor BSM interactions

- Example: charged scalar resonance with $M_S=1$ TeV, $g_l=0.3$, $g_q=0.1$

(e.g. 2nd Higgs doublet with non-standard couplings)

- LHC would measure M_S and $g_q/g_l = [BR(ud)/BR(ev)]^{1/2}$
- n decay sensitive to $G_F \cdot \epsilon_S = g_l g_q / M_S^2$ ($\epsilon_S = 1.8 \cdot 10^{-3}$)

Decay Correlations

- A: electron asymmetry
 - Perkeo II, Perkeo III, UCNA
- B: neutrino asymmetry
 - Perkeo II
- C: proton asymmetry
 - Perkeo II
- D: triple correlation
 - TRINE, emiT
- a: electron-neutrino correlation
 - aSpect, aCORN, Nab

$$\vec{\sigma}_n \cdot \vec{p}_e$$

$$\vec{\sigma}_n \cdot \vec{p}_\nu$$

$$\vec{\sigma}_n \cdot \vec{p}_p$$

$$\vec{\sigma}_n \cdot (\vec{p}_e \times \vec{p}_\nu)$$

$$\vec{p}_e \cdot \vec{p}_\nu$$

Plus Fierz interference b , helicity correlations, etc.

Principle of the *A*-coefficient Measurement

$$dW = [1 + \beta P A \cos \theta] d\Gamma(E)$$

$$A_{\text{exp}}(E) = \frac{N_1(E) - N_2(E)}{N_1(E) + N_2(E)} \approx \langle P \rangle A \beta \langle \cos \theta \rangle$$

Systematics:

Polarization

Backgrounds

(End point energy = 782 keV)

Interlude: Ultra-Cold Neutrons (UCN)

“Man is the Measure of All Things” Protagoras, 480-411 BC

- 300 neV = Potential Energy in wall
 - = $\frac{1}{2} m_n v^2$ = $\frac{1}{2} m_n (8 \text{ m/s})^2$
 - = $m_n g h$ = $m_n g (3 \text{ m})$
 - = $\frac{h^2}{2 m_n \lambda^2}$ = $\frac{h^2}{2 m_n (50 \text{ nm})^2}$
 - = $\mu_n B$ = $\mu_n (3.5 \text{ T})$
 - = $k T$ = $k (3 \text{ mK})$
- External reflection
- Running speed
- Human scale equipment
- Ultraviolet
- 100% polarization
- Ultra-cold!
- Total external reflection allows arbitrary guides and bottles; long lifetime
- Speed implies easy timing
- Installations: centimeters to meters in size
- UCN wavelength: about $0.1 \mu\text{m}$
 - *close to visible light*
 - *mirrors for people can be mirrors for UCN*
- 100% polarization is easy to achieve (for a time)

UCN can also be essentially 100 percent polarized

Backgrounds can be reduced relative to cold neutron experiments

UCN for n beta decay measurements

- Reduced background
 - Higher ratio of decays to neutrons
 - No production source background
- 100 % polarization using magnetic fields

$$V_{^{58}\text{Ni}} = \frac{2\pi\hbar^2 \rho_n a}{m} = 335 \text{ nV} \rightarrow v = 8 \text{ m/sec}$$

$$V = mgh = 102 \text{ nV/m}$$

$$V = \vec{\mu} \bullet \vec{B} = 60 \text{ nV/T}$$

But...

Production of UCNs is very tricky

Previously existing sources used the low energy tails of thermal neutron spectra from nuclear reactors, sometimes moderated to a low temperature. But this method entails high backgrounds from the reactor in any experiment.

Principle of the *A*-coefficient Measurement

$$dW = [1 + \beta P A \cos \theta] d\Gamma(E)$$

$$A_{\text{exp}}(E) = \frac{N_1(E) - N_2(E)}{N_1(E) + N_2(E)} \approx \langle P \rangle A \beta \langle \cos \theta \rangle$$

Systematics:

Polarization

Backgrounds

(End point energy = 782 keV)

UCNA Experiment — Apparatus

The UCNA Apparatus at LANL

UCNA Apparatus in LANSCE Area B

UCNA Experiment — General Approach

Novel features: UCN from pulsed spallation source
MWPC + plastic scintillator as β detector
Ultimate Goal: 0.2% measurement of A ($\delta A/A = 0.2\%$)

- Neutron Polarization
 - UCN (can produce >99% polarization with 7T magnetic field)
 - Copper and diamond-like carbon coated neutron guide (low depolarization)
- Background
 - Pulsed UCN source
 - MWPC+Plastic scintillator
- Electron backscattering
 - MWPC+Plastic scintillator
- Fiducial volume selection
 - MWPC

UCNA Experiment Detectors

- Reduced backscatters
 - Magnetic pinch
 - Low Z thin detectors
- MWPC and scintillator in coincidence
 - Position-dependent efficiency
 - Apertures and fiducial volume
- Backscatter reconstruction

In situ polarization/polarimetry

- Flipper tuned and monitored using iron foil monitor
- Polarization measured after each ~ 50 minute run
- In normal running, depolarized UCN trapped, accumulate for storage time about 30s
- Equilibrium depolarized UCN population measured by
 - change switcher to clean out right spin neutrons (not trapped)
 - change spin-flipper state to unload depol neutrons -- expect sensitivity $< 0.1\%$
 - Reloading correction necessary (UCN in guide during spin-flip)

Loading the Decay Volume: SD_2 source facility

- 2005: no decays
replaced horizontal guides w/ SS
- 2006: 2 s^{-1} (raw rate)
current: $<1\mu\text{A} \rightarrow 2\mu\text{A}$, improved flapper
- 2007: 6 s^{-1}
current: $\rightarrow 4\mu\text{A}$, source volume 2l
- 2008: 15 s^{-1}
guides changed to DLC-coated EP Cu, geom C
- 2009: 30 s^{-1}
new target and improved guides before AFP
- 2010: 55 s^{-1} (Nov)

Only source of extracted
UCN in the US

$\approx 2 \text{ UCN/cm}^3$ in decay trap!

At present Area B is best place in
world for UCNA

Solid deuterium source

UCNA Status

- Taking production data since 2009 using LANL UCN source
- Latest result: $A = -0.11954(112)$

Mendenhall *et al.*, arXiv 1210.7048 (2012)

A Correlation history

PDG λ
Derived primarily
from A [2]

0.20%

PDG 2012: $\lambda = -1.2701 \pm 0.0025$, error scaled by 1.9
From PDG2012 τ : $\lambda = -1.2748 \pm 0.0010$, error scaled by 1.8
agree at 1.75 σ level

Neutron Lifetime affects BBN

Light elements from ^2H up to ^7Li created in "first three minutes"

Weak reactions between particles:

At time ≈ 1 s, at "freeze-out" temperature $T_f \approx 1$ MeV:

neutron to proton ratio frozen to:

$$n/p = \exp(-\Delta m/kT_f) \approx 1/6.$$

$$\Delta m = m_n - m_p = 1.3 \text{ MeV}$$

After another ≈ 150 s, practically all neutron wind up in ^4He ,

i.e., He mass fraction $Y_p = 2 \times \text{neutron mass fraction} \approx 25\%$.

Primordial nucleosynthesis:

Dubbers 2013

And ratio of He/n depends directly on n lifetime: 1% lifetime uncertainty shifts calculated He fraction of $Y=0.2480 \pm 0.0003$ by 0.0015, or 5 sigma!

Free Neutron Lifetime

Precision improves over time, but accuracy?

PDG 2010: 885.7 ± 0.8 s
PDG 2011: 881.0 ± 1.5 s
PDG 2012: 880.1 ± 1.1 s

Serebrov et al.,
 $(878.5 \pm 0.7 \pm 0.3)$ seconds

Reanalysis of bottle experiments by
Serebrov, et al.

Lifetime uncertainty has grown recently

And central value has shifted by almost 1%!

PDG 2001-2010 $\tau_n = (885.7 \pm 0.8) \text{ s } (S=1)$
 Serebrov *et al.* 2005 $\tau_n = (878.5 \pm 0.8) \text{ s}$
 Pichlmaier *et al.* 2010 $\tau_n = (880.7 \pm 1.8) \text{ s}$
 PDG 2012 $\tau_n = (880.1 \pm 1.1) \text{ s } (S=1.8)$

Beam Method: Count the dying

PHYSICAL REVIEW C **71**, 055502 (2005)

FIG. 2. Experimental method for measuring lifetime by counting neutrons and trapped protons.

Danger: absolute monitor efficiency needed!

Serebrov *et al.*,
 PL B **605**, 72 (2005)

Pichlmaier *et al.*,
 PL B **693**, 221 (2010)

Bottle method: count the survivors

$$1/\tau_{\text{storage}} = 1/\tau_n + 1/\tau_{\text{loss}}$$

Wietfeldt and Greene,
 Rev. Mod. Phys. **83**, 1173 (2011)

Material bottle experiments involved 100 s extrapolations due to wall losses

Penelope experiment under development

- Superconducting multipole
 - Field zero in center eliminated by inner conductors
- Filled with UCN from FRM-2 through gap in bottom
- Decay products detected at top, guided by field lines
- Spectrum cleaned using absorber lowered from top
- Magnet now under construction

An outstanding problem: phase space evolution

- Neutron losses on scale of neutron lifetime (quasibound orbits)
- Detector efficiency changes with time
- Must fill phase space evenly, quickly: chaos!

Symmetric Trap has stable orbits Asymmetric Trap has chaotic orbits

Thin stochastic regions bounded by KAM curves.

Strong perturbation, KAM curves are destroyed.

C.-Y. Liu, 2012

UCNTau experiment designed to overcome phase space issues

Asymmetric Trap $\rightarrow$ Phase Space Mixing

- **Low symmetry** (together with **field ripples**) induces states mixing between circular orbits, through chaotic motion (or not).
- $\rightarrow$ **quick cleaning** ($\sim$ seconds) of the quasi-bound UCN with large tangential velocities.

Halbach array

- Each magnet 90° out of phase with its neighbor.
- The array has B field “ripples” of scale $L/4 = 0.5$ in
- Rotating field is orthogonal to holding field B_0

R&D: Vanadium solid state detector

Neutron scattering lengths and cross sections							
Isotope	conc	Coh b	Inc b	Coh xs	Inc xs	Scatt xs	Abs xs
V	---	-0.3824	---	0.0184	5.08	5.1	5.08
50V	0.25	7.6	---	7.3(1.1)	0.5	7.8(1.0)	60.(40.)
51V	99.75	-0.402	6.35	0.0203	5.07	5.09	4.9

Negative UCN pote Good UCN absorber

- 1.● Insert vanadium foil to absorb neutrons
- 2.● Extract foil into shielded counter
- 3.● Perform β - γ coincidence measurement.

β : 1.073 MeV , γ : 1.4 MeV

$T^{1/2} = 3.743 \text{ m}$

Use several detectors to get position information on UCN distribution inside the trap.

UCN Tracking

$$E_{\text{trap}} < E < E_{\text{trap}} + 6 \text{ neV}$$

Statistical Sensitivity

- Effective trap volume is 0.6 m^3
- UCN may trap up to $E = 48 \text{ neV}$
- LANSCE source can produce $>10 \text{ UCN cm}^{-3}$ at gate valve
- Trap density $>1 \text{ UCN/cc}$: 600,000 UCN per fill
- 15 minutes to 1 hour per fill
- 1 s precision in 6 hours

Ultimate goal: a 0.1 s total uncertainty

- Short term goals (2012-2013):
 - Prototype experiment being set up at LANSCE
 - 6 hours of DAQ for 1 s statistics
 - Measure trap lifetime
 - Detector R&D (V and monitor detectors)
 - Investigate systematic effects:
 - Spectrum dependence
 - Loss mechanisms (spin flips, residual gas, ...?)
 - Polarization and rate
 - Detector effects: position, dead time, gain
 - Workshop to discuss path to 0.1 s
- Longer term goals (2014 and out):
 - 0.1 s precision
 - 1 month of statistics at present UCN production rate

UCNTau Experiment now commissioning

- Asymmetric magnet
- Holding
- UCNs from through
- Neutron on vanadium trap, active
- Has storage

B History

VALUE
0.9807 ± 0.0030 OUR AVERAGE

0.9802 ± 0.0034 ± 0.0036

0.967 ± 0.006 ± 0.010

0.9801 ± 0.0046

0.9894 ± 0.0083

1.00 ± 0.05

0.995 ± 0.034

DOCUMENT ID TECN COMMENT

SCHUMANN 07 CNTR Cold *n*, polarized

KREUZ 05 CNTR Cold *n*, polarized

SEREBROV 98 CNTR Cold *n*, polarized

KUZNETSOV 95 CNTR Cold *n*, polarized

CHRISTENSEN70 CNTR Cold *n*, polarized

EROZOLIM... 70c CNTR Cold *n*, polarized

P =

0.997(1)

0.987(5)

0.975(3)

0.669(2)

0.87(3)

0.74(1)

PERKEO IIB

M. Schumann et al.,
PRL 99, 191803 (2007);
and Ph.D. Thesis

$$B_{\text{exp}}(E_e) = \frac{N^{--}(E_e) - N^{++}(E_e)}{N^{--}(E_e) + N^{++}(E_e)}$$

$B_{\text{exp}}(E_e)$ related to B via
integration of W over the
hemispheres

The "UCNB" concept

[W.S. Wilburn et al.,
submitted to Rev. Mex. Fis.]

1 Tesla, B-Field

Basic Idea: Detect e-p
coincidences

e: "start" (prompt)

p: "slow" ($\sim \mu\text{s} - \text{ms}$)

UCNB Observables

Neutron Polarization

Electron and Proton Directions

Electron Energy

First-Pass Goal:
0.1% on B

Highly Segmented Si Detector

- Developed by LANL with LDRD funding
- .5, 1, 1.5 and 2mm thick
- New level of resolution
- Background reduction with e-p coincidence
- <100nm dead layer, p-gun

Preliminary Systematics + Statistical Error

Table 1: Major corrections and error table for “Bexp (E)”			
Major Systematic	Correction	Error	Source
Statistics	n/a	0.05%	Assuming a year long run
Neutron polarization	<0.1%	0.02%	UCNA experience
Det. calibration, linearity, response	0.10%	<0.01%	Projected silicon detector linearity
Proton backscatter	0.07%	<0.01%	[8]
Electron Backscatter	0.30%	0.06%	[10]
Missed coincidences	0.10%	<0.03%	1 ms window [9]
Accidental coincidences	-0.10%	<0.01%	Rate 10/s ~ 1 ms window [9]
Mirroring effects in B field expansion	-0.04%	0.02%	UCN absorbers
Total	0.30%	<0.07%	“Cexp (E)”

Electron Backscatter Modeled

PHYSICAL REVIEW C **68**, 055503 (2003)

- Monte Carlos exist – UCNA
- Only effects energy in “C”
- Can use bias to under stand low energy response

125 KeV data from Caltech e-gun

UCN Hall Layout

Neutron Polarizing/Spin-flipper Magnet

UCN Source

Superconducting Spectrometer

UCNA -> UCNb, UCNB

Current facility DOE and NSF funded

UCNA->UCNE

UCNb

UCN sources, Solid, LHe, Doppler

Classic Doppler Source

1-2 ucn/cc

New “Superthermal” UCN sources utilize phonon scattering. Both liquid and solid media are being used.

What makes a good superthermal UCN source?

- Low neutron absorption
- High single phonon energy
- Light atoms
- Weak crystal

LANSCCE LINAC provides unique, highly-flexible beam delivery to multiple facilities
7 months per yr @ 24/7 with ~1200 user visits

The Los Alamos Neutron Science Center

800 MeV protons, up to 1 mA at 120 Hz and 12% DF

Lujan Center

- *Materials science and condensed matter research*
- *Bio-science*
- *Nuclear physics*
- *BES National User Facility*

WNR

- *Nuclear physics*
- *Semiconductor irradiation*

Proton Radiography

- *HE science, dynamic materials science, hydrodynamics*

Isotope Production Facility

- *Nuclear medicine*
- *Research isotope production*

Ultracold Neutrons Facility

- *Fundamental nuclear physics*

The heart of LANSCE is a very flexible 800-MeV proton linear accelerator (LINAC)--one of the most powerful in the world!

[illegible]

(Nights and weekends!)

[illegible]

Top view of experiment hall

The UCNA Apparatus at LANL

UCNA Apparatus in LANSCE Area B

UCN Source Layout

Cold neutron measurements: direct detection and Ar activation

Argon activation cold flux measurements

- Measure background
- Freeze calibrated volume of Ar into cryostat
- Exposed to known proton flux, producing ^{41}Ar
- Boil off into calibrated gas volume
- Count decays with HPGe counter
- Calibrate efficiency and solid angle with ^{60}Co source
- Compare to MCNP prediction
- Extract cold neutron flux: $(1.7 \pm 0.3) \times 10^9$ n/cm²/s/ μA (with no deuterium)

Cold neutron TOF measurement

Cold neutron TOF measurement

- Cold neutrons detected after 3.6 m vertical flight through 2 cm diameter hole in shield package
- 1 proton pulse per second, 1.4×10^{10} protons per 250 ns pulse
- Compare to MCNP prediction: 1.56×10^5 n/ μ C detected vs 1.70×10^5 predicted (full); 0.75×10^5 n/ μ C detected vs 0.86×10^5 predicted (empty) (< 25 meV)

Internal UCN Measurements

Internal UCN Measurements

- Cut off in MC corresponds to height + Al potential – D2 potential
- UCN transport using custom MC: track back to UCN density in D2 of 85(10) UCN/ $\mu\text{C}/\text{cm}^3$
- Use measured CN density + Atchison production cross sections: 107(20) UCN/ $\mu\text{C}/\text{cm}^3$
- Compare to prototype:
- 460(90) UCN/ $\mu\text{C}/\text{cm}^3$
- @6 μA : 3000 UCN/ cm^3

External UCN Measurements: Transport thru shield stack

System lifetime measurements

- Simulate transport of UCN through shield wall, detect with GV monitor, Gate valve closed
- Vary specularity and loss per bounce to match measured distributions
- ~10 cm diameter SS guides, butted joints, ~6 m total length to gate valve
- Measure lifetime with flapper open (D2 exposed) and closed (D2 hidden)
- Monitor detector port not well modeled: check normalization with main detector

Flapper Open: sensitive to specularity

Flapper Closed: sensitive to loss

External UCN Measurements: Confirmation of monitor normalization

- Main vs monitor comparison checks modeling of monitor port
- Varying magnetic field also allows extraction of UCN velocity distribution
- Calibrated MC allows prediction of UCN density throughout system
- Track back to density in D2: $60(12) \text{ UCN}/\mu\text{C}/\text{cm}^3$
- UCN density at monitor port @ $6 \mu\text{A}$: $79(16) \text{ UCN}/\text{cm}^3$

Direct Measurement of UCN density at Shield Exit

- Direct measurement of UCN density using V foil
- 1 cm diameter, 250 μm thick
- $^{51}\text{V} + n \rightarrow ^{52}\text{V} \rightarrow ^{52}\text{Cr} + \beta + 1.43 \text{ MeV } \gamma$
- Calibrate solid angle with ^{60}Co source
- Correct for thermal upscatter in V foil (1.02)
- Correct for V oxide potential (1.16)
- UCN density @ 6 μA is then 52(9) UCN/ cm^3
- Compare to 79(16) UCN/ cm^3 measured indirectly

Layout of Test Port Area

User Experiment location in LANSCE Area B

Neutron Lifetime Experiment on Test Port

The Bottom Line

- Source and Test Port are available and running now
- Parameters:
 - LANSCE runs 7 months/year
 - Proton beam is shared with PRAD
 - available ~100 hrs/wk while accelerator is on
 - UCN source is shared with UCNA
 - Test port beam can be on 10 minutes per hour while UCNA runs
 - ~25 UCN/cc at Test Port (after PPM), 52 at shield wall, up to 180 neV (at 5 kW incident proton power)
 - UCNs at Test Port are polarized to be high-field seekers
 - Backgrounds outside of beam gate are largely natural
 - Beam gate is 0.2 s per 5 s
- Allocation by UCNA Executive Committee for now
 - But we hope for a PAC process soon

Future Improvements???

- Improved proton beam tune
 - $x \sim 2$
 - Requires 0.2 M\$ for new beam pipe and diagnostics
- Beam pattern: spread out pulses
 - $x \sim 2$
 - Requires 0.5 M\$ for safety equipment
- Lower loss, higher V guides
 - $x \sim 3$
 - Requires 0.3 M\$ and six months to replace guides
- Duty factor: kick beam to pRad
 - $x \sim 2$
 - Requires 3 M\$ for kicker and shield wall

Future Improvements???

- Improved proton beam tune
 - $x \sim 2$
 - Requires 0.2 M\$ for new beam pipe and diagnostics
- Beam pattern: spread out pulses
 - $x \sim 2$
 - Requires 0.5 M\$ for safety equipment
- Lower loss, higher V guides
 - $x \sim 3$
 - Requires 0.3 M\$ and six months to replace guides
- Duty factor: kick beam to pRad
 - $x \sim 2$
 - Requires 3 M\$ for kicker and shield wall

UCN Production Improvement by Spreading out Beam pulses

- Data taken in August, 2012, at low current
- Matching pulse spacing to lifetime above flapper optimizes UCN production at equal average current
- Present spacing limited by beamline current limiting devices (hardware), not by safety requirements

Future Improvements???

- Improved proton beam tune
 - $x \sim 2$
 - Requires 0.2 M\$ for new beam pipe and diagnostics
- Beam pattern: spread out pulses
 - $x \sim 2$
 - Requires 0.5 M\$ for safety equipment
- Lower loss, higher V guides
 - $x \sim 3$
 - Requires 0.3 M\$ and six months to replace guides
- Duty factor: kick beam to pRad
 - $x \sim 2$
 - Requires 3 M\$ for kicker and shield wall
 - Wall alone increases time available to UCN by $\sim 20\%$
 - Kicker in addition increases time by 100%

UCNA Collaboration

[R. W. Pattie, Jr.](#),^{1,2} [J. Anaya](#),³ [H. O. Back](#),^{1,2} [J. G. Boissevain](#),³ [T. J. Bowles](#),³ [L. J. Broussard](#),^{2,4} [R. Carr](#),⁵ [D. J. Clark](#),³ [S. Currie](#),³ [S. Du](#),¹ [B. W. Filippone](#),⁵ [P. Geltenbort](#),⁶ [A. García](#),⁷ [A. Hawari](#),⁸ [K. P. Hickerson](#),⁵ [R. Hill](#),³ [M. Hino](#),⁹ [S. A. Hoedl](#),^{7,10} [G. E. Hogan](#),³ [A. T. Holley](#),¹ [T. M. Ito](#),^{3,5} [T. Kawai](#),⁹ [K. Kirch](#),³ [S. Kitagaki](#),¹¹ [S. K. Lamoreaux](#),³ [C.-Y. Liu](#),¹⁰ [J. Liu](#),⁵ [M. Makela](#),^{3,12} [R. R. Mammei](#),¹² [J. W. Martin](#),^{5,13} [D. Melconian](#),^{7,14} [N. Meier](#),¹ [M. P. Mendenhall](#),⁵ [C. L. Morris](#),³ [R. Mortensen](#),³ [A. Pichlmaier](#),³ [M. L. Pitt](#),¹² [B. Plaster](#),^{5,15} [J. C. Ramsey](#),³ [R. Rios](#),^{3,16} [K. Sabourov](#),¹ [A. L. Sallaska](#),⁷ [A. Saunders](#),³ [R. Schmid](#),⁵ [S. Seestrom](#),³ [C. Servicky](#),¹ [S. K. L. Sjue](#),⁷ [D. Smith](#),¹ [W. E. Sondheim](#),³ [E. Tatar](#),¹⁶ [W. Teasdale](#),³ [C. Terai](#),¹ [B. Tipton](#),⁵ [M. Utsuro](#),⁹ [R. B. Vogelaar](#),¹² [B. W. Wehring](#),⁸ [Y. P. Xu](#),¹ [A. R. Young](#),^{1,2} and [J. Yuan](#)⁵

1Department of Physics, North Carolina State University, Raleigh, North Carolina 27695, USA

2Triangle Universities Nuclear Laboratory, Durham, North Carolina 27708, USA

3Los Alamos National Laboratory, Los Alamos, New Mexico 87545, USA

4Department of Physics, Duke University, Durham, North Carolina 27708, USA

5W. K. Kellogg Radiation Laboratory, California Institute of Technology, Pasadena, California 91125, USA

6Institut Laue-Langevin, 38042 Grenoble Cedex 9, France

7Physics Department, University of Washington, Seattle, Washington 98195, USA

8Department of Nuclear Engineering, North Carolina State University, Raleigh, North Carolina 27695, USA

9Research Reactor Institute, Kyoto University, Kumatori, Osaka, 590-0401, Japan

10Physics Department, Princeton University, Princeton, New Jersey 08544, USA

11Tohoku University, Sendai 980-8578, Japan

12Department of Physics, Virginia Tech, Blacksburg, Virginia 24061, USA

13Department of Physics, University of Winnipeg, Winnipeg, MB R3B 2E9, Canada

14Cyclotron Institute, Texas A&M University, College Station, Texas 77843, USA

15Department of Physics and Astronomy, University of Kentucky, Lexington, Kentucky 40506, USA

16Department of Physics, Idaho State University, Pocatello, Idaho 83209, USA

UCN τ and UCNB Collaborations

“UCN τ ” Collaboration

S. Clayton, M. Makela, C. Morris, J. Ramsey, A. Saunders, P. Walstrom, Z. Wang (LANL)

D. Bowman, S. Penttila (ORNL)

E. Adamek, C. Cude, W. Fox, M. Hozo, C.-Y. Liu, N. Callahan, D. Salvat, J. Vanderwerp, B. Slaughter, K. Solberg (Indiana)

B. Vogelaar (VT)

K. Hickerson (Cal Tech)

A. R. Young, E. Dees (NCSU)

“UCNB” Collaboration

R. Alarcon, S. Balascuta, L. Broussard, R. Carr, V. Cirigliano, B. Filippone,
A. García, P. Geltenbort, D. Harris, A. Hawari, K. Hickerson, S. Hoedl,
A. Holley, G. Hogan, J. Liu, M. Makela, R. Mammei, J. Martin,
P. McGaughey, D. Melconian, M. Mendenhall, C. Morris, R. Pattie,
A. Pérez-Galván, M. Pitt, B. Plaster, J. Ramsey, R. Rios, A. Salas-Bacci,
A. Saunders, S. Seestrom, E. Tatar, R.B. Vogelaar, B. Wehring,
W.S. Wilburn, A. Young

Conclusions

- Full suite of neutron beta decay experiments at Los Alamos
- UCNs give unique advantages
- UCNA producing competitive results
 - Will eventually reach $\sim 0.2\%$
- UCN τ will start acquiring data soon
- UCNB will probe new territory
- LANSCE UCN Source is in production
 - Supplying beam to UCNA, nEDM, UCN τ , UCNb
- Source performance has been measured and modeled
 - CN and UCN measurements match MC predictions
- Incremental improvements are still possible
- Technology is not near fundamental limits